Astro 507 Lecture 10 Feb. 12, 2020

Announcements:

• Preflight 2 posted, due noon Friday

includes discussion question on the Anthropic Principle!

 Office Hours: Instructor after class today, or by appt TA: noon- 1pm tomorrow, or by appt

Last time:

 $\vdash$ 

- $\bullet$  completed cosmic inventory:  $\Omega_0 \approx 1$  and  $\Omega_{matter} \approx 0.3~Q$ : and so?
- high time to become relativistic introduced invariant interval  $\Delta s^2 \ Q$ : wut? why? particle with mass m, relativistic energy E has speed:  $v(E) = \sqrt{1 - m^2/E^2} \ Q$ : consequences?

### **Causality and Spacetime**

any particle of total energy E, mass m

moves at speed  $v(E) = \sqrt{1 - \left(\frac{m}{E}\right)^2}$ 

- massive particles have  $0 \le v < c$
- massless particles (e.g.,  $\gamma)$  have v=c
- $\Rightarrow v = c = 1$  is universal speed limit
- ⇒ cannot transmit particles, info any faster

future light cone at spacetime point p
encloses region within which
particles/info can move

i.e., region p can influence

- $\Rightarrow$  future light cone is spacetime region
- $_{\scriptscriptstyle \rm N}$   $\,$  causally connected to p

*past light cone* at *p Q*: *significance*?



#### past light cone at p

events in cone can send particles/info to pi.e., region which could have influenced p $\Rightarrow$  past light cone=causally connected to p



ω

Q: two events causally connected if?

*Q:* sufficient or just necessary?

# What About Gravity?

A. Einstein (1905):

Newtonian dynamics  $\rightarrow$  relativistic dynamics space, time  $\rightarrow$  spacetime forever more

Relativity and classical fields:

- E&M: Maxwell eqs relativistically OK! (*motivated* Lorentz, SR)
- Newtonian gravity:  $\vec{g} = -\nabla \phi = -Gm/r^2 \hat{r}$  and  $\nabla^2 \phi = 4\pi G\rho$ an *unmitigated disaster Q: Why?*

How to fix?

4

First attempt: analogy with electrostatics Q: why?

$$\nabla^2 \phi - \partial_t^2 \phi = 4\pi G\rho \tag{1}$$

- bad news: disagrees with expt (gives no light bending)
- good news: right "flavor" e.g., operator  $\nabla^2 - \partial_t^2 \rightarrow$  waves  $\rightarrow$  gravitational radiation!

### **Mystic Pisa**

**Experiment: Galileo** (Tower of Pisa?) free fall independent of mass, size, shape, composition *Q: lawyer's fine print?* 

**Theory: Newton** always:  $\vec{F} = m\vec{a}$ gravity: mass is "coupling strength"  $\Rightarrow \vec{F}_{grav} = m\vec{g}$  $\Rightarrow$  free fall has  $\vec{a} = \vec{g} \rightarrow$  indep of object properties interesting curiosity

#### **Theory: Einstein**

ហ

gravity is acceleration, so maybe acceleration is gravity

i.e., their physical effects indistinguishable/equivalent

### **Equivalence Principle**

T-shirt summary (R. Wald):

all bodies fall the same way in a gravitational field

an observer in free fall Q: meaning?

cannot perform any experiment to determine whether she is in a gravitational field

#### an observer undergoing acceleration

cannot perform any experiment to determine whether she is in a gravity field or an accelerating spacecraft

*Q:* explain apple weight–Earth's surface

б

vs rocket accelerating a = g?

*Q*: explain apple drop–Earth's surface vs rocket with a = g?

## Newton's Apple Experiment: Two Views

Isaac Newton on Earth's surface:

- holds an apple in his hand, **pushes up with force** F = mg says: must oppose weight so net force zero
- releases apple, observes downward acceleration says: motion due to net gravity force

Albert Einstein in rocket with constant acceleration a = g:

- holds apple in hand, pushes up with force F = mg says: to keep apple in my non-inertial accelerating hand must push so it accelerates too
- releases apple, observes downward acceleration says: motion due to my non-inertial frame

Note: identical physical results, radically different explanations

1

Q: what about horizontal light beam?

*Q*: what about horizontal ball toss?

## **Gravity Bends Light**

Rocket Experiment: www: illuminating animation
in accelerating rocket, shoot a horizontal beam
\* light ray deflected
\* entire light path bent (in fact, a parabola!)
"gravity's rainbow"

But by equivalence principle: must find same result due to gravity, so: \* gravity bends light rays gravitational lensing



 $\infty$ 

## **Acceleration and Photons**

Still consider accelerating spaceship

Experiment: light signals between top & bottom each flashes every  $\tau_{\rm em} = 1$  sec according to emitter's clock emission frequency  $t_{\rm em} = 1/\tau_{\rm em}$ 

Q: what is change in top clocks' speed when pulse arrives? Q: what frequency does top clock observe? asymmetry: top clock accelerates away from bottom flash  $\rightarrow$  relative speed changes during light transit by amount  $\delta v_{top} = -a\delta t \simeq -ah/c$ sign  $\rightarrow$  receding from source

 $\rightarrow$  top observer sees freq Doppler shifted downward: *redshift* 

$$f_{\rm obs,top} \approx \left(1 - \frac{\delta v}{c}\right) f_{\rm em,bottom}$$
 (2)

so top observer sees bottom flash period as

10

$$\frac{\tau_{\rm obs} - \tau_{\rm em}}{\tau_{\rm em}} = \frac{\delta\tau}{\tau} = -\frac{\delta f}{f} \approx \frac{\delta v}{c} = +\frac{ah}{c^2}$$
(3)

Q: which means? and upon applying equivalence principle...?

### **Equivalence Principle and Photon Properties**

Equivalence principle: gravitational results identical to rocket

- shifted frequencies: gravitational redshift/blueshift!
- period shift: gravitational time dilation

$$\frac{\delta t}{t} = \frac{\delta \lambda}{\lambda} \approx \frac{gh}{c^2} = \frac{\phi}{c^2} \tag{4}$$

attic clocks faster than basement clocks: verified experimentally! www: Pound-Rebka expt

in weak gravity:

- fractional shift  $\approx \phi/c^2$
- $\bullet$  set by change in gravitational potential  $\phi$

 $\frac{1}{1}$ 

## **Consequences of the Equivalence Principle**

equivalence principle implies that gravity

- bends light trajectory: *distorts path in space*
- changes apparent frequency: *distorts apparent flow of time*

together these mean  $\rightarrow$  gravity alters spacetime!

Einstein (1915): include gravity in spacetime

## **General Relativity**

Newton (1687): Universal Gravitation
gravity is a force (field) that couples to mass
matter tells gravity how to force

gravity force tells matter how to move

Einstein (1915): General Relativity

gravity is spacetime curvature: not a force!

- $\star$  "matter tells spacetime how to curve
- $\star$  spacetime tells matter how to move" –J. .A. Wheeler

### Curved Spacetime?

Curved space: geometric constructions in space (triangles, rectangles, circles... *Q: how define?*) give non-Euclidean results *Q: namely?* 

13

*Q: so-curved spacetime?* 

### **Spacetime Curvature**

Test: (Feynman Lectures II, Chapter 42)

- construct geometric object in spacetime
- are properties Euclidean?

Case 1: Minkowski Space (i.e., special relativity, no accel) (1-D) interval ("line element") for events separated by (dt, dx)

$$ds^2 = dt^2 - dx^2 \tag{5}$$

Construct rhombus: in *spacetime* two observers go from events A to B $\triangleright$  obs 1: go left at v = 0.5c for 10 s, then wait 10 s  $\triangleright$  obs 2: wait 10 s, then go left at v = 0.5c for 10 s

14

Q: spacetime diagram?



result is Euclidean Q: why?

 $\Rightarrow$  Minkowski spacetime is **not curved = flat** 

Case 2: Surface of Earth (i.e., const accel: gravity) (1-D) line element:

$$ds^{2} = \left(1 + \frac{2\phi}{c^{2}}\right) dt^{2} - \left(1 + \frac{2\phi}{c^{2}}\right)^{-1} dx^{2}$$
(6)

where  $\phi = \phi(x)$ : time-independent Newtonian potential

Construct rhombus in *spacetime* two observers go from events *A* to *B* ▷ obs 1: go up 1 km, then wait 10 s ▷ obs 2: wait 10 s, then go up 1 km

Q: spacetime diagram?



result is **not Euclidean**:

(wait time) = 
$$(\delta s)_{\text{wait}} = \sqrt{1 + 2gh/c^2} \ (\delta t)_{\text{wait}}$$
 (7)

why? waiting time "advance differently" - time dilation!

 $\stackrel{\prec}{\neg} \Rightarrow \text{Earth's spacetime is } \frac{}{\text{gravity}} \Rightarrow \frac{}{\text{spacetime curvature}}$ 

# **GR** on a **T-Shirt**

General Relativity spirit and approach: like special relativity, only moreso

Special Relativity concepts retained:

- **spacetime**: events, relationships among them
- interval gives observer-independent (invariant) measure of "distance" between events
- Special Relativity is a special case of GR
   SR: no gravity → no curvature → "flat spacetime"
   GR limit: gravity sources→0 give spacetime→Minkowski

GR: Special Relativity concepts generalized

- gravity encoded in spacetime structure
- spacetime can be curved
- coordinates have no intrinsic meaning

18

## **The Metric**

Fundamental object in GR: metric

consider two nearby events, separated by coordinate differences  $dx = (dx^0, dx^1, dx^2, dx^3)$  GR (in orthogonal spacetimes) sez:

interval between them given by "line element"

$$ds^{2} = A(x) (dx^{0})^{2} - B(x) (dx^{1})^{2} - C(x) (dx^{2})^{2} - D(x) (dx^{3})^{2}$$
  
$$\equiv \sum_{\mu\nu} g_{\mu\nu} dx^{\mu} dx^{\nu} \equiv g_{\mu\nu} dx^{\mu} dx^{\nu}$$

where the **metric tensor**  $g_{\mu\nu}$ 

- in this case (orthogonal spacetime): g = diag(A, B, C, D)
- components generally are functions of space & time coords
- is symmetric, i.e.,  $g_{\mu\nu} = g_{\nu\mu}$
- encodes all physics (like wavefunction in QM)
   *Q*: *if no gravity=Minkowski, what's the metric*?

physical interpretation of interval: like in SR

$$ds^2$$
 = (apparent elapsed time)<sup>2</sup>  
- (apparent spatial separation)<sup>2</sup>

$$dx = event 2$$

$$dt \int ds$$

$$event 1$$

t

★ observers have *timelike* worldlines:  $ds^2 > 0$ ★ light has *null* ds = 0 worldlines

Simplest example: Minkowski space (Special Relativity)  $g_{\mu\nu} = \text{diag}(1, -1, -1, -1)$ : constant values proper spatial distances:

- i.e., results using meter sticks
- measured simultaneously  $(dx^0 = 0)$

length element:

21

 $d\ell^2 = -ds^2 = d\ell_1^2 + d\ell_2^2 + d\ell_3^2 = g_{11}(dx^1)^2 + g_{22}(dx^2)^2 + g_{33}(dx^3)^2$ <br/>space (3-)volume element:

$$dV_3 = d\ell_1 d\ell_2 d\ell_3 = \sqrt{|g_{11}g_{22}g_{33}|} dx^1 dx^2 dx^3$$

spacetime 4-volume element:

$$dV_4 = d\ell_0 dV_3 = \sqrt{|g_{00}g_{11}g_{22}g_{33}|} dx^0 dx^1 dx^2 dx^3$$
$$= \sqrt{|\det g|} dx^0 dx^1 dx^2 dx^3$$

Example: Minkowski space, Cartesian coords

$$ds^{2} = dt^{2} - dx^{2} - dy^{2} - dz^{2}$$

length:  $d\ell^2 = dx^2 + dy^2 + dz^2$ 3-volume:  $dV_3 = dx \, dy \, dz$ 4-volume:  $dV_4 = dx \, dy \, dz \, dt$ 

Example: Minkowski space, spherical coords

$$ds^{2} = dt^{2} - dr^{2} - r^{2}d\theta^{2} - r^{2}\sin^{2}\theta d\phi^{2}$$
  
length:  $d\ell^{2} = dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2})$   
3-volume:  $dV_{3} = r^{2}\sin\theta dr d\theta d\phi \equiv r^{2}dr d\Omega$   
4-volume:  $dV_{4} = r^{2}dr d\Omega dt$ 

22



### **Cosmological Spacetimes**

Want to describe spacetime of the universe to zeroth order: homogeneous, isotropic

at each spacetime point
 exactly one observer sees isotropy\*
 call these fundamental observers
 roughly: "galaxies" i.e., us
 (strictly speaking, we don't qualify) Q: why?

2. isotropy at each point  $\rightarrow$  homogeneity but can be homogeneous & not isotropic

24

\*We will see: observers moving w.r.t. FOs eventually come to rest w.r.t. FOs

### 3. homogeneity and isotropy $\rightarrow$ symmetries

### U. is "maximally symmetric"

- $\rightarrow$  greatly constrain allowed spacetimes
  - i.e., allowed metrics

## **The Cosmic Line Element**

cosmological principle: can divide spacetime into time "slices" i.e., 3-D spatial (hyper) surfaces

> populated by fundamental observers

```
\triangleright with coords, e.g., (t, x, y, z)
```

 $\triangleright$  choose FO's to have  $d\vec{x} = 0$ 

i.e., spatial coords are **comoving** ("fixed to expanding grid") on surface: fundamental observers must all have  $ds^2 = dt^2 \rightarrow i.e., g_{tt} = const = 1 \ Q$ : why?  $\rightarrow g_{tt}$  indep of space, time

 $_{N}$  these give:

$$ds^2 = dt^2 - g_{ii}(dx^i)^2$$
 (8)

## **Cosmological Principle and the Cosmic Metric**

#### homogeneity and time

no space dependence on  $d\ell_0 = dt$ 

- can define cosmic time t (FO clocks)
- at fixed t, time lapse dt not "warped" across space

#### homogeneity and space

- at any *t*, properties invariant under translations
- no center
- can pick arbitrary point to be origin
- e.g., here!

Cosmological spacetime encoded via cosmic **metric** which determines how the interval depends on coordinates any observer computes interval between events as  $ds^2 = (elapsed time)^2 - (spatial displacement)^2$ 

Cosmic metric so far:

$$ds^2 = dt^2 - g_{ii}(dx^i)^2$$
(9)

where: t is cosmic time

now impose *isotropy* 

200

- at any cosmic t, interval invariant under rotations
- pick arbitrary origin, then (comoving) spherical coords the usual  $r, \theta, \phi$ , with  $r^2 = x^2 + y^2 + z^2$  and arbitrary origin (usually, but not always, here!)

For *fundamental* observers, maximal symmetry demands metric which  $can^*$  be written as:

$$ds^{2} = dt^{2} - a(t)^{2} d\ell_{\text{com}}^{2}$$
(10)  
=  $dt^{2} - a(t)^{2} \left[ f(r) dr^{2} + r^{2} (d\theta^{2} + \sin^{2} \theta d\phi^{2}) \right]$ (11)

a(t) is the cosmic scale factor f(r) is as yet undetermined

- for flat (Euclidean) space, f(r) = 1
- so  $f \neq 1 \rightarrow$  non-Euclidean spatial geometry = curved space!

Q: why same time dep for radial and angular displacements? Note power of cosmo principle

 $\rightarrow$  only allowed dynamics is uniform expansion a(t)!

\*other space & time coordinates possible and sometimes useful but in all cases space and time must *factor* in this way

20